Impact of ozone cross-section choice on (WF)DOAS total ozone retrieval

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DOAS total ozone retrieval and ozone temperature

Weighting function DOAS

$$\ln \frac{I_{obs}}{F_{obs}} = \ln \left(\frac{I}{F}\right)_{mod} + \frac{d \ln(I/F)}{d \operatorname{Toz}} \mid_{mod} (\operatorname{TOZ}_{fit} - \operatorname{TOZ}_{clim}) + \frac{d \ln(I/F)}{d \operatorname{T}} \mid_{mod} (\operatorname{T}_{fit} - \operatorname{T}_{clim}) + \dots + \operatorname{Pol}$$

Coldewey-Egbers et al., 2005

Radiation transfer model

"Standard" DOAS

$$\ln \frac{I_{obs}}{F_{obs}} = -SC_1 \cdot \sigma(T_1) - SC_2 \cdot \sigma(T_2) + \\ + \cdots + Pol \qquad T_1 \approx 220 \text{ K} \\ T_2 \approx 240 \text{ K} \\ TOZ = (SC_1 + SC_2) / AMF$$
Roozendael et al., 2006
Radiation transfer model

- ➔ DOAS satellite retrievals
- 325-335 nm (WFDOAS: 336.6-335 nm)
- Weighting function DOAS (Coldewey-Egbers et al., 2005, Weber et al., 2005)
 - Scalar temperature shift in the a-priori temperature profile
- Standard DOAS: ESA operational retrieval (Roozendael et al. 2006)





Ozone and temperature terms in WFDOAS equation

Anti-correlation between ozone and ozone temperature term

Depending on fitting window size and position correlation ranges between r = -0.4 and -0.6

Coldewey-Egbers et al., 2005



WFDOAS total ozone data sets/cross-section used



- WFDOAS applied to GOME (since 1995), SCIAMACHY (since 2002), and GOME2 (since 2006)
 - →GOME1/ERS : Burrows et al. 1999 (GOME FM), shift: +0.17 nm
 - →SCIAMACHY/ENVISAT Bogumil et al., 2003 (SCIA FM), scaled 5.3%, shift: +0.08 nm
 - →GOME2/METOP A: Burrows et al., 1999, convolved, shift: +0.017nm



GOME WFDOAS comparison to ground (at Hohenpeissenberg)

- Very good agreement, higher seasonality in the differences with Dobsons
 - Error sources:
 - temperature dependence of cross-section (fixed T in ground retrievals)
 - Dobson error: 1.3%/10K (Komhyr et al. 1988)
 - OBrewer error: 0.0-0.9%/10K (Kerr 2002)
 - ➔ Different x-sections
 - Bass-Paur (Brew/Dobs)
 vs. Burrows et al.
 (GOME)
 - → Instrumental issues
 Ostraylight issue at high SZA (all instruments)
 - Calibration and optical degradation of satellite instruments (UV sensors)



Hohenpeissenberg 48 °N,11 °E



Cross-section data tested

Ozone cross-section data \rightarrow GOME FM (Burrows et al., 1999) OT=202 K, 221 K, 241 K, 273 K, 293 K Ospectral resolution: 0.17 nm @ 330 nm → SCIAMACHY FM (Bogumil et al., 2003) OT=203 K, 223 K, 243 K, 273 K, 293 K Ospectral resolution: 0.20 nm @ 330 nm → GOME2 FM3 V3 (Gür et al. 2005) OT=(203 K), 223 K, 243 K, 273 K, 293 K currently under revision Ospectral resolution: 0.29 nm @ 330nm → Bass-Paur (Paur and Bass, 1985) OT=203 K, 218 K, 228 K, 243 K, 273 K, 298 K Ospectral resolution: ~0.05 nm(?) \rightarrow Malicet et al. (1995), Brion et al., (1993), Daumont et al. (1998)OT=218 K, 228 K, 243 K, 273 K, 295 K Ospectral resolution: 0.01-0.02 nm



Temperature parametrerisation after Bass-Paur (t temperature) in $^{\circ}C$, λ wavelength in nm):

$$\sigma(\lambda, t) = a_0(\lambda) \left[1 + a_1(\lambda)t + a_2(\lambda)t^2 \right]$$

Direct comparisons of cross-sections/Solar fitting

Slit function derived from comparison between (a) cross-sections and (b) from solar data

$$\sigma'(\lambda, T) = c_0 \cdot \sigma(\lambda + \Delta\lambda, T) \otimes r(\lambda) + Pol(\lambda)$$
$$I'(\lambda) = c_0 \cdot I(\lambda + \Delta\lambda) \otimes r(\lambda) + Pol(\lambda)$$



- → Scaling factor: c_o
- → wavelength shift: $\Im \lambda$
- \rightarrow Instrument function r(λ)
- → Polynomial: $Pol(\lambda)$
- High spectral resolution reference data:
 - → Kitt-Peak FTS solar data
 - ➔ Brion O3 cross-sections



8

Cross-section issues in satellite DOAS I

Cross-sections	T[K]	Shift [nm]	Scaling [-]	Gaussian FWHM [nm]	Solar FWHM [nm]	Scaling wrt. GOME FM
Burrows et al. 1999 GOME FM	225 240	+0.017(2) +0.017(2)	1.027(2) 1.023(2)	0.158(5) 0.159(6)	0.165(8)	-
Bogumil et al. 2003	225	+0.008(2)	0.970(3)	0.222(6)	0.202(13)	-5.6%
SCIAMACHY FM	240	+0.009(2)	0.973(3)	0.219(6)		-4.9%
Bass & Paur, 1985	225	+0.020(3)	1.015(3)	0.079(10)	-	+1.2%
(original, MPI database)	240	+0.023(5)	1.000(3)	0.087(10)		+2.3%
Bass & Paur, 1985	225	+0.024(4)	1.004(3)	0.084(12)	-	+2.3%
(ACSO web page)	240	+0.021(4)	1.004(3)	0.078(11)		+1.9%

Slitfunction fits w.r.t. Brion cross-sections: 326.6-334.5 nm

- Shifts minimising fit residuals in WFDOAS agree with shifts from cross-section comparisons (good wavelength calibration for Brion et al.)
- SCIAMACHY FM requires a scaling of +5% wrt to GOME1
- Note: total ozone change is -6DU/0.01nm shift (or -2%/0.01 nm) !!!



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Retrieval test: SCIAMACHY & GOME2



- Brion et al. & Bass-Paur (ACSO) ozone agree well and are 2-3% higher than SCIA/GOME-FM
- Using proper scaling for SCIA-FM, agreement with GOME-FM below 0.5%



retrieval test/ ozone temperature change



 Small changes in tO3 differences from SCIA to GOME related to changes in ozone temperature

Retrieval test/ spectral fit rms



- higher fit residuals with Bass-Paur
- Brion et al. similar (SCIA) or slightly lower (GOME2) than SCIA/GOME-FM



Summary and conclusion

- Comparison of different cross-sections used in satellite retrievals require
 Proper wavelength shifts

 Odirect comparison of cross-sections
 Overification by minimisation of spectral fit residuals

 Proper scaling

 Odirect comparison of cross-sections
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 Odirect comparison of cross-sections
 Overification of scaling by retrieval comparisons
 Osome compensation/changes can occur due to interference of ozone and temperature terms in the (WF)DOAS equation
- Satellite DOAS ozone retrievals in the 325-335 nm window
 - ➔ Brion cross-sections are well wavelength calibrated (no shifts required)
 - ➔ Bass-Paur (ACSO) with shifts of 0.022 nm produces same results as Brion et al. but with higher spectral fit residuals
 - ➔ GOME-FM and scaled SCIA-FM (+5%) provide same ozone results, but use of Brion/Bass-Paur leads to differences of +2 to +3% in retrieved ozone
 - → Fit residuals from Brion and GOME/SCIA-FMs are similar

Requirements on "new" UV/Vis x-section measurements

- Important requirements for laboratory measurements
 - → Wavelength calibration (-2% O3 error/0.01nm shift)
 - → Absolute scaling, e.g. cross-path experiment with 253.65 nm as reference
 - Temperature stability & sufficient numbers of temperature points
 - → Good slit function characterisation and/or sufficient high spectral resolution (~0.01 nm)
- ESA HARMONICS study
 - new cross-section measurements (talk by Anna Serdyuchenko)
 - ➔ Reanalysis of GOME, SCIA, and GOME2 x-section data

